## On Symmetries of Linear and Nonlinear PDEs

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- Point Symmetries of PDEs
- Infinite Symmetries of Linear PDEs
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- Point Symmetry Structure of Linear PDEs
- 5 A Nonlinear Example a 2D Hyperelastic Model

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## Point Symmetries of PDEs

**2** Infinite Symmetries of Linear PDEs

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## Point Symmetries of PDEs

### Model equations

• A given system of PDEs of order k:

$$R^{\sigma}[\mathbf{u}] = R^{\sigma}(\mathbf{x}, \mathbf{u}, \partial \mathbf{u}, \dots, \partial^{k}\mathbf{u}) = 0, \quad \sigma = 1, \dots, N.$$

• Higher-order symmetries are rather uncommon (though important). We will talk about point symmetries in this lecture.

#### Point symmetries

• A one-parameter Lie group of point transformations preserving the model:

$$\begin{aligned} (x^{i})^{*} &= f^{i}(\mathbf{x}, \mathbf{u}; \varepsilon) = x^{i} + \varepsilon \xi^{i}(\mathbf{x}, \mathbf{u}) + O(\varepsilon^{2}), \\ (u^{\mu})^{*} &= g^{\mu}(\mathbf{x}, \mathbf{u}; \varepsilon) = u^{\mu} + \varepsilon \eta^{\mu}(\mathbf{x}, \mathbf{u}) + O(\varepsilon^{2}). \end{aligned}$$

• The corresponding infinitesimal generator (tangent vector field):

$$\mathbf{X} = \xi^{i}(\mathbf{x}, \mathbf{u}) \frac{\partial}{\partial x^{i}} + \eta^{\mu}(\mathbf{x}, \mathbf{u}) \frac{\partial}{\partial u^{\mu}}.$$

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- Finite-dimensional group general case for nonlinear PDEs.
- Infinite-dimensional group (parameterized by arbitrary function(s) with of fewer arguments than # of independent variables) occurs for nonlinear models (e.g., Galilei group).
- Infinite-dimensional group (parameterized by arbitrary function(s), # arguments = # independent variables) - common for linear PDEs, and PDEs that may be linearized by a point transformation.

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## Point Symmetries of PDEs

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• If the given PDE system

$$R^{\sigma}[\mathbf{u}] = R^{\sigma}(\mathbf{x}, \mathbf{u}, \partial \mathbf{u}, \dots, \partial^{k}\mathbf{u}) = 0, \quad \sigma = 1, \dots, N$$

is linear, then the evolutionary symmetry components are arbitrary solutions of the linearized equations (linear homogeneous PDEs)

$$\mathcal{L}\{R\}^{\sigma}_{\mu}[\mathbf{u}]\zeta^{\mu}=\mathbf{0},\quad \sigma=1,\ldots,N.$$

• Example: u = u(x, t),

$$u_t = u_{xx},$$

has an infinite point symmetry group with  $X = g(x, t) \frac{\partial}{\partial u}$ ,

$$x^* = x$$
,  $t^* = t$ ,  $u^* = u + g(x, t)$ ,

where  $g_t = g_{xx}$ .

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## Infinite symmetries - linearization by a point transformation





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**Theorem 2.4.1** (Necessary conditions for the existence of an invertible linearization mapping of a nonlinear PDE system). If there exists an invertible mapping  $\mu$  of a given nonlinear PDE system  $\mathbf{R}\{x;u\}$  ( $m \ge 2$ ) to some linear PDE system  $\mathbf{S}\{z;w\}$ , then

(i)  $\mu$  is a point transformation of the form

$$z^{j} = \phi^{j}(x, u), \quad j = 1, ..., n,$$
 (2.61a)

$$w^{\gamma} = \psi^{\gamma}(x, u), \quad \gamma = 1, ..., m;$$
 (2.61b)

(ii) R{x; u} has an infinite set of point symmetries given by an infinitesimal generator

$$\mathbf{X} = \xi^{i}(x, u)\frac{\partial}{\partial x^{i}} + \eta^{\nu}(x, u)\frac{\partial}{\partial u^{\nu}}$$
(2.62)

with infinitesimals  $\xi_i(x, u)$ ,  $\eta^{\nu}(x, u)$  of the form

$$\xi^{i}(x, u) = \alpha^{i}_{\sigma}(x, u)F^{\sigma}(x, u), \qquad (2.63a)$$

$$\eta^{\nu}(x, u) = \beta^{\nu}_{\sigma}(x, u)F^{\sigma}(x, u),$$
 (2.63b)

where  $\alpha_{\sigma}^{i}(x, u), \beta_{\sigma}^{\nu}(x, u), i = 1, ..., n; \nu; \sigma = 1, ..., m$ , are specific functions of x and u, and where  $F = (F^{1}, ..., F^{m})$  is an arbitrary solution of some linear PDE system

$$L[X]F = 0$$
 (2.64)

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in terms of some linear differential operator L[X] and specific independent variables  $X = (X^1(x, u), \dots, X^n(x, u)) = (\phi^1, \dots, \phi^n).$ 

## Infinite symmetries - linearization by a point transformation

- Some PDE systems can be linearized by a nonlocal transformation (have a sufficiently large set of nonlocal symmetries).
- E.g. Burgers equation:  $u_t + uu_x u_{xx} = 0$ , u = u(x, t): finitely many point/contact symmetries.
- Potential equations  $v_x = 2u$ ,  $v_t = 2u_x u^2$  have an infinite number of point symmetries given by the infinitesimal generator

$$\mathbf{X} = e^{\nu/4} \left\{ [2h(x,t) + g(x,t)u] \frac{\partial}{\partial u} + 4g(x,t) \frac{\partial}{\partial v} \right\},\$$

where (g(x, t), h(x, t)) is an arbitrary solution of the linear PDE system

$$h=g_x, \quad h_x=g_t.$$

• As a result, the Hopf-Cole transformation  $u = 2y_x/y$  maps (non-invertibly) the Burgers equation into a linear diffusion equation:

$$\frac{\partial}{\partial x}\left(\frac{2}{y}(y_t-y_{xx})\right)=0.$$

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• We will mostly follow this paper:



# Cheviakov, A. (2010).

Symbolic computation of local symmetries of nonlinear and linear partial and ordinary differential equations. *Mathematics in Computer Science*, 4(2-3), 203–222.

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• Consider a linear DE system

$$L^{\sigma}(x, u, \partial u, \dots, \partial^{k} u) = F^{\sigma}(x), \quad \sigma = 1, \dots, N,$$

of order k, with  $n \ge 2$  independent variables  $x = (x^1, ..., x^n)$ , and  $m \ge 1$  dependent variables  $u(x) = (u^1(x), ..., u^m(x))$ .

- Here each  $L^{\sigma}[u]$  is a linear homogeneous differential expression in u(x).
- If u(x) is a solution of the linear system L<sup>σ</sup>[u] = F<sup>σ</sup>(x), and w(x) is a solution of the linear homogeneous system L<sup>σ</sup>[w] = 0, then

$$\hat{u}(x) = u(x) + w(x)$$

is also a solution of the linear system:  $L^{\sigma}[\hat{u}] = F^{\sigma}(x)$ .

• The transformation  $u(x) \rightarrow u(x) + w(x)$  yields an infinite set of *trivial* Lie point symmetries

$$\mathbf{X}_{tr} = w^{\mu}(x) \frac{\partial}{\partial u^{\mu}}.$$

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### Linear ODEs

- For linear ODEs, the dimension of Lie group is always finite, which is better from the point of view of symbolic computations.
- If the explicit form of the general solution of the linear homogeneous ODE is unknown, symbolic software would not be able to compute all point symmetries explicitly.

### Linear PDEs

- For linear PDEs, the dimension of point symmetry Lie algebra is infinity.
- Trivial symmetry components are general solutions of linear homogeneous PDE(s).
- These PDEs do not have closed-form solutions, so symbolic software is not able to compute these symmetries explicitly.

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• The following helpful theorems has been established in this paper:



### Bluman, G. (1990).

Simplifying the form of Lie groups admitted by a given differential equation. Journal of mathematical analysis and applications, 145(1), 52-62.

#### Theorem

Suppose L[u] = F(x) is a scalar linear PDE (i.e.,  $N = m = 1, n \ge 2$ ) of order  $k \ge 2$ . Then components  $\xi^i, \eta$  of its point symmetries satisfy

$$\frac{\partial \xi^i}{\partial u} = \frac{\partial^2 \eta}{\partial u^2} = 0, \quad i = 1, \dots, n.$$

#### Theorem

Suppose L[u] = F(x) is a scalar linear ODE (i.e., N = n = n = 1) of order  $k \ge 3$ . Then components  $\xi, \eta$  of its point symmetries satisfy

$$\frac{\partial\xi}{\partial u} = \frac{\partial^2\eta}{\partial u^2} = 0.$$

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Ovsiannikov, L. V. (1982).

Group Analysis of Differential Equations, Academic Press.

#### Ovsiannikov's "linear DE conjecture"

For a linear DE system  $L^{\sigma}[\mathbf{u}] = F^{\sigma}(\mathbf{x})$ , components  $\xi^{i}, \eta^{\mu}$  of any point symmetry

$$\mathbf{X} = \xi^{i}(\mathbf{x}, \mathbf{u}) \frac{\partial}{\partial x^{i}} + \eta^{\mu}(\mathbf{x}, \mathbf{u}) \frac{\partial}{\partial u^{\mu}}$$

satisfy

$$\frac{\partial \xi^{i}}{\partial u^{\nu}} = 0, \quad \frac{\partial^{2} \eta^{\mu}}{\partial u^{\nu} u^{\lambda}} = 0, \quad i = 1, \dots, n, \quad \mu, \nu, \lambda = 1, \dots, m.$$

It is stated to hold for the "majority of linear DEs" (that is, PDE and ODE systems).Is it true?

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#### Ovsiannikov's "linear DE conjecture"

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satisfy

$$\frac{\partial \xi^{i}}{\partial u^{\nu}} = 0, \quad \frac{\partial^{2} \eta^{\mu}}{\partial u^{\nu} u^{\lambda}} = 0, \quad i = 1, \dots, n, \quad \mu, \nu, \lambda = 1, \dots, m.$$

• The conjecture can be verified symbolically: use <> capability of rifsimp.

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- **()** Symmetries of the linear homogeneous heat equation:  $u_t = u_{xx}$ .
- **3** Symmetries of the linear non-homogeneous heat equation:  $u_t = u_{xx} + f(x)$ .
- **③** Symmetries of the linear wave equation:  $u_{tt}$

$$u_{tt}=c^2(x)u_{xx}.$$

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Fig. 1. Material and Eulerian coordinates.

### Material picture

- A solid body occupies the reference (Lagrangian) volume  $\Omega_0 \subset \mathbb{R}^3$ .
- Actual (Eulerian) configuration:  $\Omega \subset \mathbb{R}^3$ .
- Material points are labelled by  $\mathbf{X} \in \Omega_0$ .
- The actual position of a material point:  $\mathbf{x} = \phi(\mathbf{X}, t) \in \Omega$ .

Image: A math a math



Fig. 1. Material and Eulerian coordinates.

### Material picture

- Velocity of a material point X:  $\mathbf{v}(\mathbf{X}, t) = \frac{d\mathbf{x}}{dt}$ .
- Jacobian matrix (deformation gradient):

$$\mathbf{F}(\mathbf{X}, t) = \nabla \phi; \quad J = \det \mathbf{F} > 0;$$
$$\mathbf{F} = \{F_{ii}\} = \{F_{ii}^{i}\}.$$



Fig. 1. Material and Eulerian coordinates.

### Material picture

- Boundary force (per unit area) in Eulerian configuration:  $t = \sigma n$ .
- Boundary force (per unit area) in Lagrangian configuration: T = PN.
- $\sigma = \sigma(\mathbf{x}, t)$  is the Cauchy stress tensor.
- $\mathbf{P} = J \boldsymbol{\sigma} \mathbf{F}^{-T}$  is the first Piola-Kirchhoff tensor.



Fig. 1. Material and Eulerian coordinates.

## Material picture

- Density in reference configuration:  $\rho_0 = \rho_0(\mathbf{X})$  (time-independent).
- Density in actual configuration:

$$\rho = \rho(\mathbf{X}, t) = \rho_0/J.$$

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### Equations of motion (no dissipation, purely elastic setting):

 $\rho_0 \mathbf{x}_{tt} = \mathsf{div}_{(X)} \mathbf{P} + \rho_0 \mathbf{R},$ 

•  $\mathbf{R} = \mathbf{R}(\mathbf{X}, t)$ : total body force per unit mass.

•  $(\operatorname{div}_{(X)}\mathbf{P})^i = \frac{\partial P^{ij}}{\partial X^j}.$ 

Cauchy stress tensor symmetry (conservation of angular momentum):

$$\mathbf{F}\mathbf{P}^{\mathsf{T}} = \mathbf{P}\mathbf{F}^{\mathsf{T}} \quad \Leftrightarrow \quad \boldsymbol{\sigma} = \boldsymbol{\sigma}^{\mathsf{T}}.$$

The first Piola-Kirchhoff stress tensor:

$$\mathbf{P} = 
ho_0 \frac{\partial W}{\partial \mathbf{F}}, \qquad P^{ij} = 
ho_0 \frac{\partial W}{\partial F_{ii}}.$$

•  $W = W(\mathbf{X}, \mathbf{F})$ : a scalar strain energy density function.

A. Cheviakov (UofS, Canada)

Symmetries of DEs

• Strain energy density W depends only on certain matrix invariants:

$$W=U(I_1,I_2,I_3)=\overline{U}(\overline{I}_1,\overline{I}_2,\overline{I}_3).$$

• For the left Cauchy-Green strain tensor  $\mathbf{B} = \mathbf{F}\mathbf{F}^{\mathsf{T}}$ ,

$$\begin{split} I_1 &= \operatorname{Tr} \mathbf{B} = F^i{}_k F^i{}_k = \lambda_1^2 + \lambda_2^2 + \lambda_3^2, \\ I_2 &= \frac{1}{2} [(\operatorname{Tr} \mathbf{B})^2 - \operatorname{Tr}(\mathbf{B}^2)] = \lambda_1^2 \lambda_2^2 + \lambda_2^2 \lambda_3^2 + \lambda_1^2 \lambda_3^2, , \\ I_3 &= \det \mathbf{B} = J^2 = \lambda_1^2 \lambda_2^2 \lambda_3^2. \end{split}$$

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• Strain energy density W depends only on certain matrix invariants:

$$W = U(I_1, I_2, I_3) = \overline{U}(\overline{I}_1, \overline{I}_2, \overline{I}_3)$$

#### Table 1: Neo-Hookean and Mooney-Rivlin constitutive models

Туре	Neo-Hookean	Mooney-Rivlin
Standard	$W = aI_1, a > 0.$	$W = aI_1 + bI_2,$ a, b > 0
Generalized	$W = a\bar{h}_1 + c(J-1)^2,$ a, c > 0.	$W = a\overline{h}_1 + b\overline{h}_2 + c(J-1)^2$ a, b, c > 0
Generalized (Ciarlet) "compressible"	$W = a h_1 + \Gamma(J),$ $\Gamma(q) = c q^2 - d \log q,  a, c, d > 0$	$W = al_1 + bl_2 + \Gamma(J)$ $\Gamma(q) = cq^2 - d\log q,  a, b, c, d > 0$

Image: A math a math

• Strain energy density W depends only on certain matrix invariants:

$$W = U(I_1, I_2, I_3) = \overline{U}(\overline{I}_1, \overline{I}_2, \overline{I}_3).$$

#### Example: the Neo-Hookean Case

- Strain energy density:  $W = a I_1$ , a = const.
- Equations of motion are linear and decoupled:

$$(x^{k})_{tt} = a \left( \frac{\partial^2}{\partial (X^1)^2} + \frac{\partial^2}{\partial (X^2)^2} + \frac{\partial^2}{\partial (X^3)^2} \right) x^{k},$$
  
k = 1, 2, 3.

• Strain energy density W depends only on certain matrix invariants:

$$W = U(I_1, I_2, I_3) = \overline{U}(\overline{I}_1, \overline{I}_2, \overline{I}_3).$$

### General compressible framework

• Strain energy density:

$$W = \overline{U}(\overline{I}_1, \overline{I}_2, \overline{I}_3).$$

Barred invariants:

$$\bar{I}_1 = J^{-2/3} I_1, \quad \bar{I}_2 = J^{-4/3} I_2, \quad \bar{I}_3 = J$$

• Piola-Kirchhoff stress tensor - incompressible case:

$$\mathbf{P} = -\boldsymbol{p} \, \mathbf{F}^{-T} + \rho_0 \frac{\partial W}{\partial \mathbf{F}}, \qquad \boldsymbol{P}^{ij} = -\boldsymbol{p} \, (\boldsymbol{F}^{-1})^{ji} + \rho_0 \frac{\partial W}{\partial F_{ij}}.$$

Image: A math a math

# 2D incompressible Mooney-Rivlin hyperelastic materials

• A 2D setting:

$$x^{1,2} = x^{1,2} \left( X^1, X^2, t \right), \qquad x^3 = X^3,$$
  
 $R^{1,2,3} = 0, \qquad \rho_0 = \text{const.}$ 

• Derivative notation:

$$\frac{\partial^2 x^1}{\partial t^2} \equiv x_{tt}^1, \quad \frac{\partial x^1}{\partial X^2} \equiv x_2^1, \quad \frac{\partial^2 x^2}{\partial X^1 \partial X^2} \equiv x_{12}^2, \quad \frac{\partial^2 p}{\partial X^2 \partial t} \equiv p_{2t},$$

• Equations of motion:

$$R^{1}[x, p] = 1 - J = 1 - \left(x_{1}^{1}x_{2}^{2} - x_{2}^{1}x_{1}^{2}\right) = 0,$$
  

$$R^{2}[x, p] = x_{tt}^{1} - \left[\alpha\left(x_{11}^{1} + x_{22}^{1}\right) - p_{1}x_{2}^{2} + p_{2}x_{1}^{2}\right] = 0,$$
  

$$R^{3}[x, p] = x_{tt}^{2} - \left[\alpha\left(x_{11}^{2} + x_{22}^{2}\right) - p_{2}x_{1}^{1} + p_{1}x_{2}^{1}\right] = 0,$$

where  $\alpha = 2(a + b) = \text{const} > 0$  is a material parameter.

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• The model has a classical Lagrangian:

$$\mathcal{L} = -K + W - p(J-1),$$

where

$$K = rac{1}{2} \left( (x_t^1)^2 + (x_t^2)^2 
ight)$$

is the kinetic energy density,

$$W = \frac{\alpha}{2} \left( \left( x_1^1 \right)^2 + \left( x_2^1 \right)^2 + \left( x_1^2 \right)^2 + \left( x_2^2 \right)^2 \right)$$

is the and potential (strain) energy density.

• A Lagrangian density equivalent to the above (up to a total divergence) can be obtained using a homotopy formula

$$\widehat{\mathcal{L}} = \int_0^1 u \cdot R[\lambda u] \, d\lambda.$$

The Mooney-Rivlin equations arise as Euler-Lagrange equations under the actions of the Euler operators:

$$\mathbf{E}_p \, \mathcal{L} = R^1[x, p], \qquad \mathbf{E}_{x^1} \, \mathcal{L} = R^2[x, p], \qquad \mathbf{E}_{x^2} \, \mathcal{L} = R^3[x, p].$$

## Extended Kovalevskaya form

• The above 2D equations admit an extended Kovalevskaya form:

$$\begin{split} \widehat{R}^{1}[x,p] &= x_{1}^{1} - \left(x_{2}^{2}\right)^{-1} \mathcal{S}[x^{1},x^{2}] = 0, \\ \widehat{R}^{2}[x,p] &= x_{11}^{2} - \left(-x_{22}^{2} + \alpha^{-1} \left[x_{tt}^{2} - p_{1}x_{2}^{1} + p_{2}\left(x_{2}^{2}\right)^{-1} \mathcal{S}[x^{1},x^{2}]\right]\right) = 0, \\ \widehat{R}^{3}[x,p] &= p_{1} - M[x^{1},x^{2}] \left\{ \left(x_{2}^{2}\right)^{2} \left(x_{2}^{1}x_{tt}^{2} - x_{2}^{2}x_{tt}^{1}\right) + \left(x_{1}^{2}\mathcal{N}[x^{1},x^{2}] + x_{2}^{1}\right)x_{2}^{2}p_{2} \right. \\ &\left. + \alpha \left[x_{2}^{2}x_{22}^{1}\mathcal{N}[x^{1},x^{2}] - x_{2}^{2}x_{12}^{2} - \left(x_{2}^{1}\mathcal{N}[x^{1},x^{2}] + x_{1}^{2}\right)x_{2}^{2}\right] \right\} = 0, \end{split}$$

where

$$\begin{split} \mathcal{N}[x^1, x^2] &= \left(x_2^1\right)^2 + \left(x_2^2\right)^2, \quad \mathcal{M}[x^1, x^2] = \mathcal{N}[x^1, x^2]^{-1} \left(x_2^2\right)^{-2}, \\ \mathcal{S}[x^1, x^2] &= \left(1 + x_2^1 x_1^2\right). \end{split}$$

• The leading derivatives:  $\{x_1^1, x_{11}^2, p_1\}$ .

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#### Theorem

The two-dimensional Mooney-Rivlin equations are invariant under an infinite-dimensional group of Lie point transformations given by the infinitesimal generators

$$\begin{split} \mathrm{R}^{1} &= \frac{\partial}{\partial t} , \quad \mathrm{R}^{2} = \frac{\partial}{\partial X^{1}} , \quad \mathrm{R}^{3} = \frac{\partial}{\partial X^{2}} , \\ \mathrm{R}^{4} &= X^{2} \frac{\partial}{\partial X^{1}} - X^{1} \frac{\partial}{\partial X^{2}} , \quad \mathrm{R}^{5} = x^{2} \frac{\partial}{\partial x^{1}} - x^{1} \frac{\partial}{\partial x^{2}} , \\ \mathrm{R}^{6} &= F_{1}(t) \frac{\partial}{\partial x^{1}} - F_{1}^{\prime\prime}(t) x^{1} \frac{\partial}{\partial p} , \\ \mathrm{R}^{7} &= F_{2}(t) \frac{\partial}{\partial x^{2}} - F_{2}^{\prime\prime}(t) x^{2} \frac{\partial}{\partial p} , \\ \mathrm{R}^{8} &= F_{3}(t) \frac{\partial}{\partial p} , \\ \mathrm{R}^{9} &= t \frac{\partial}{\partial t} + X^{1} \frac{\partial}{\partial X^{1}} + X^{2} \frac{\partial}{\partial X^{2}} + x^{1} \frac{\partial}{\partial x^{1}} + x^{2} \frac{\partial}{\partial x^{2}} , \end{split}$$

where  $F_1(t)$ ,  $F_2(t)$ , and  $F_3(t)$  are arbitrary functions of time.

• Maple/GeM computation

• Evolutionary forms and Noether theorem: next slide.

## Cheviakov, A. and St. Jean, S. (2015).

A comparison of conservation law construction approaches for the two-dimensional incompressible Mooney-Rivlin hyperelasticity model. *JMP* 56, 121505.

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